

Chile *r* forecasting: How to improve & validate *r*-forecasts

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S4 Collab Meeting summer 2024

Bullet Points Handed to Us:

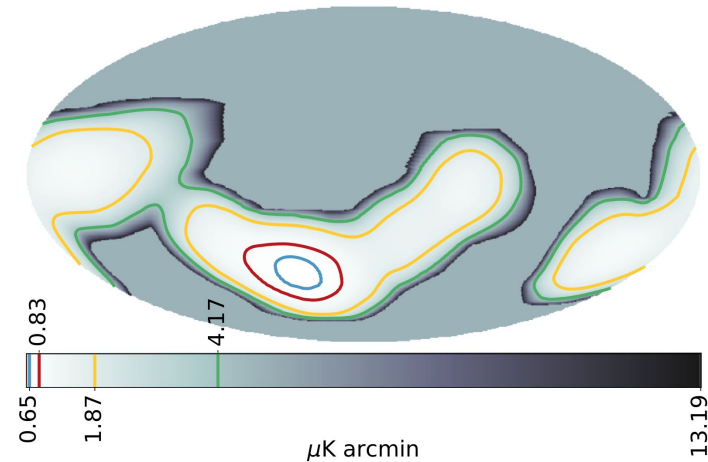
1. De-lensing validation
2. Map based validation
3. Optimizing band allocation on SATs and delensing LATs

4. Inclusion of potential systematics

Question / comment: validation (of Alt 3) vs. optimization (of Chilean survey) ?

Current Status: 1) Delensing validation (map based)

- ❖ As [Raphael said yesterday](#) we have seen recovery of lensing B-mode templates from Julien Carron and Sebastian Belkner that match predictions pretty well.
- ❖ The current DC11 round of sims includes all-Chile alt3.
 - Lensing reconstruction of that is still in progress.
 - How to include the extended alt3 field with its highly variable delensing level into re-analyses will take some work but is probably not a huge issue...
 - FG cleaning of LAT maps complete (Shamik)



Current Status: 2) Map-based validation

- ❖ Fisher calcs yield $\sigma(r)$ only - map based sims can give bias on r as well
- ❖ To back the AoA Fisher studies we have done latest DC10/11 map based runs which include alt1, alt2 and alt3 (all-Chile)
- ❖ Colin gave a good [summary of the current results](#) yesterday - there is significant bias for the medium and complex models - very strong for alt3 (all-Chile) if one just uses the less good sky regions blindly.

Question: Should we worry about method-based variations?

Question: Should we take (more) seriously high complexity FG?

3) (Re)Optimizing band allocation on SATs and delensing LATs

- ❖ The SAT/LAT split and band-allocations within each were (supposed to have been) optimized in the past - [see published paper](#)
- ❖ This should maybe be revisited - particularly for the all-Chile configs...

Question: Specifically, do we need HF and LF bands for the delensing survey?
How does the answer depend on sky complexity?

Question: Do current FG models span the right range of complexity for delensing?

Question: Do extragalactic FG matter for delensing?

Other questions ?

Big unknowns

- ❖ Polarized atmosphere?
- ❖ Foreground bias?
- ❖ Systematics?

Anything else (before we discuss systematics at length)?

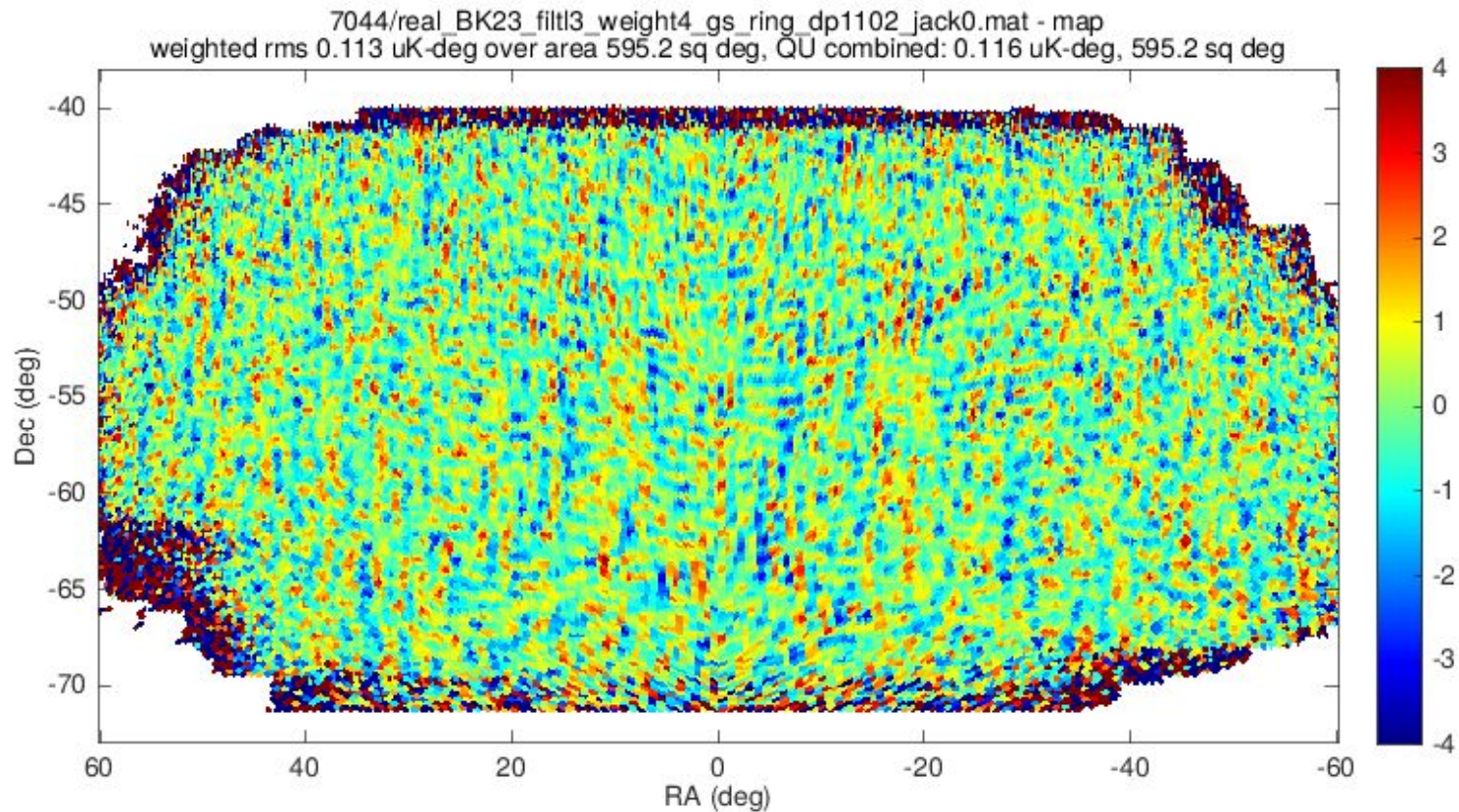


4) Inclusion of potential systematics: Switching to “aggressive optics”

- ❖ In the existing calcs the SAT map sensitivity is scaled from BK achieved performance using the relative number of detectors and calculated NET ratios.
 - This is simple to change for an “aggressive optics” version.
- ❖ Way more uncertain is whether “aggressive optics” will be OK from a bias on r point of view
 - All we know for sure is that such bias has been controlled to a level of a few times the noise down to current BK noise levels. (i.e. map null tests pass)
 - And as we see below the systematic effects already being removed like scan-sync are very large compared to the signal we are after ($r=0.001$)
 - Knowingly making such systematics bigger seems a brave move...

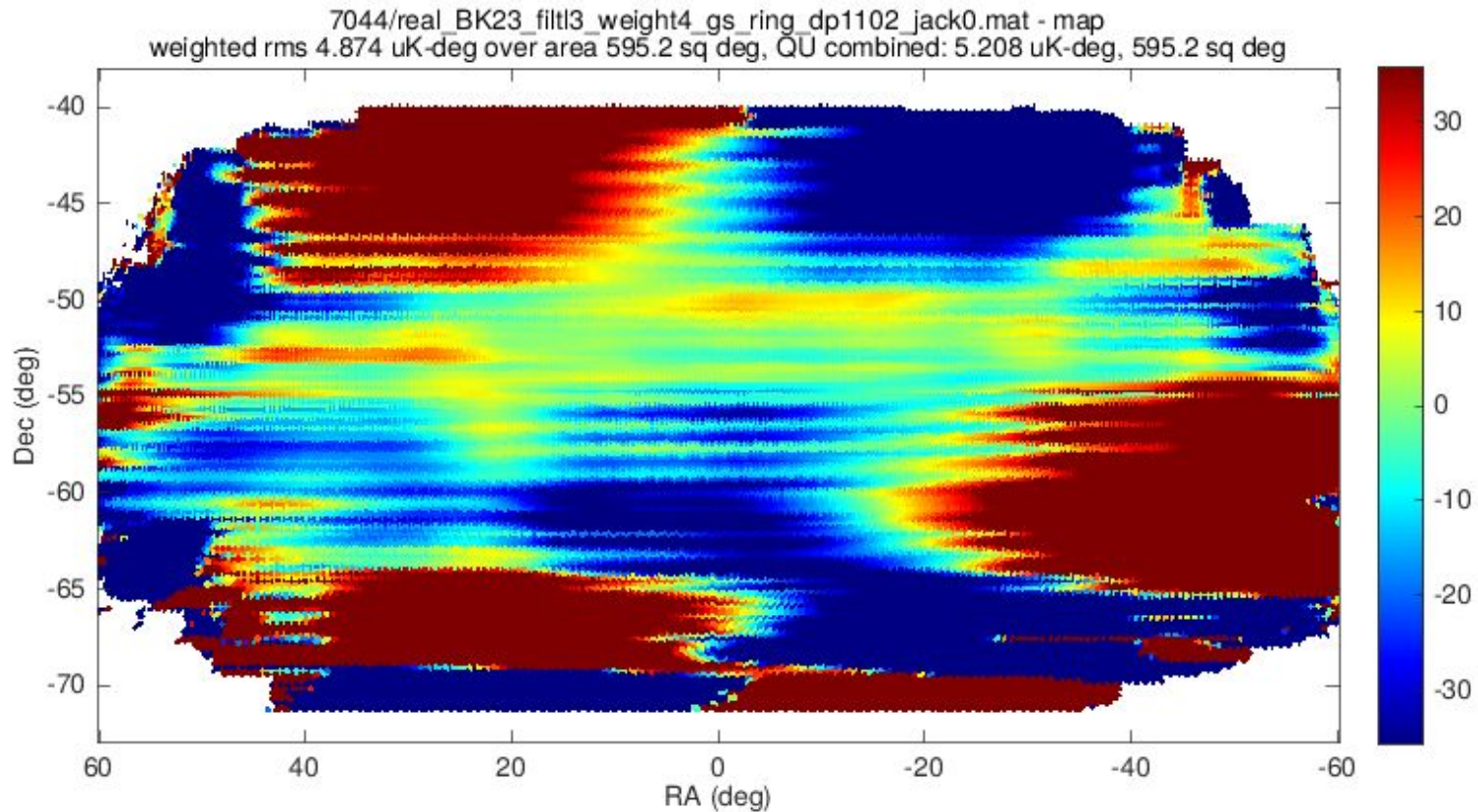
The BICEP3 8-year Q map

The remaining map after filtering (LCDM E-mode dominated)



The BICEP3 8-year Q map

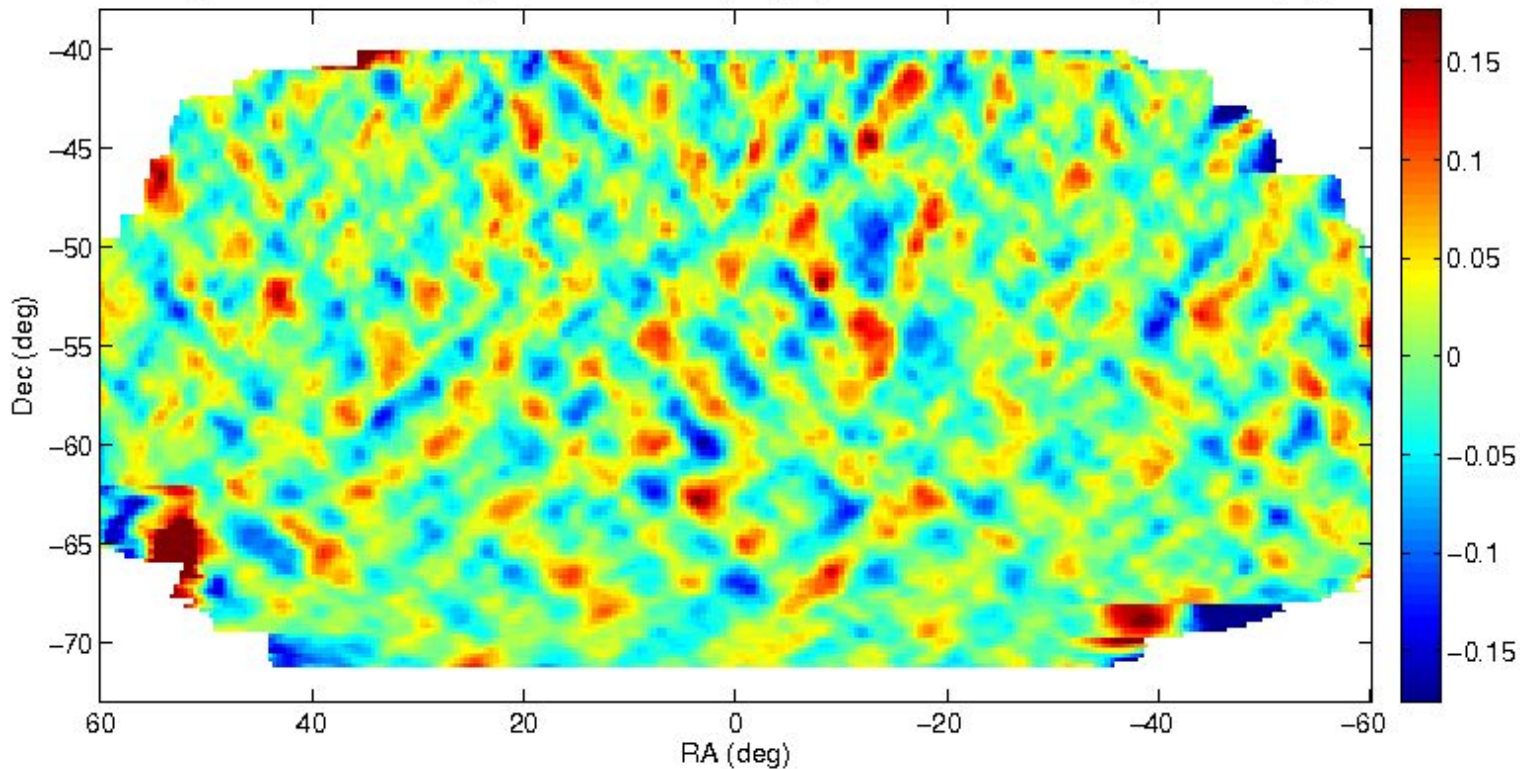
The part removed by scan-synch filter (note change in color scale)



The BICEP3 8-year Q map

An $r=0.1$ signal (note change in color scale)

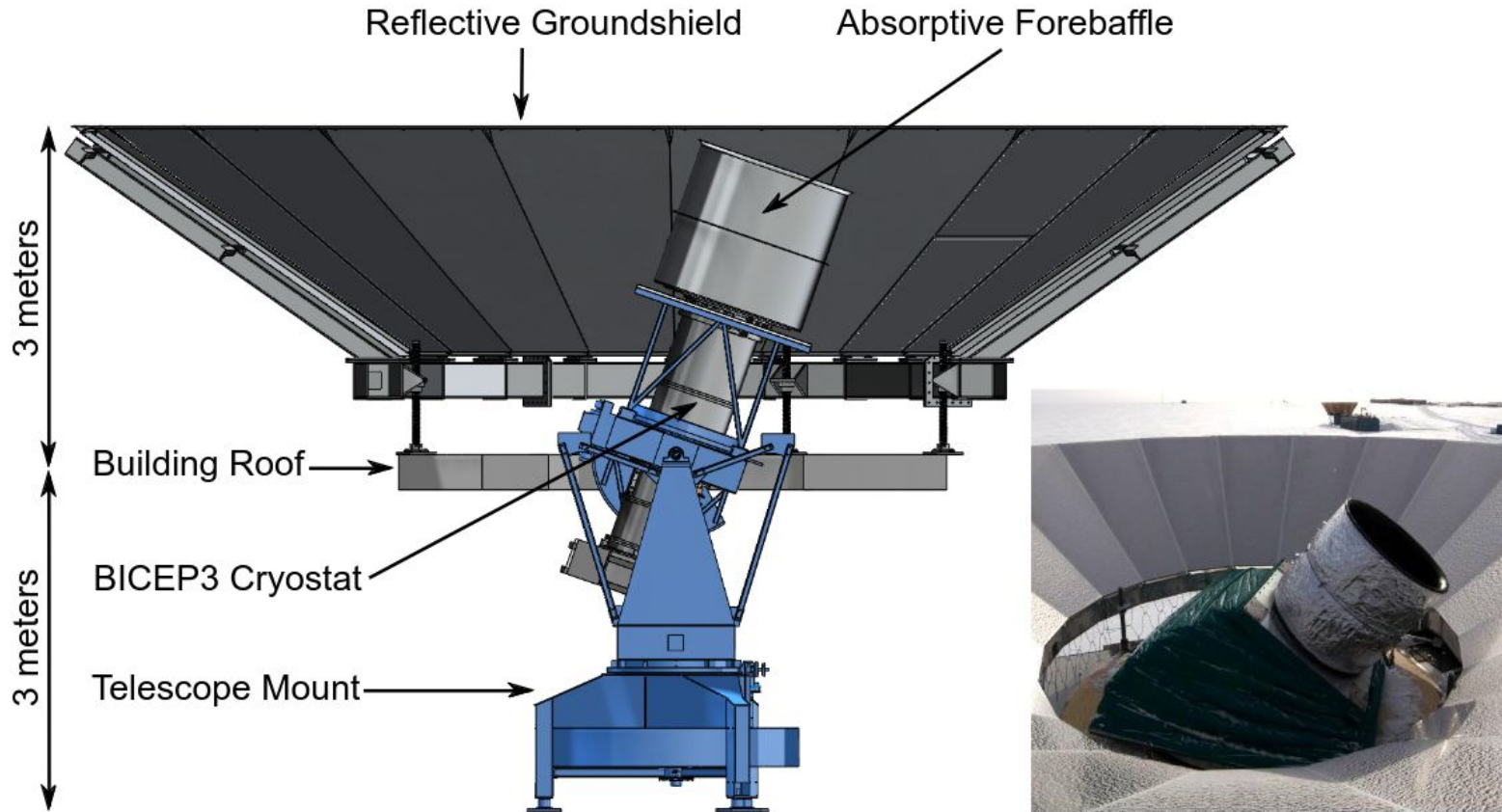
`r_filtp3_weight3_gs_dp1100_jack0.mat x 5260/0014_aabdefgh_filtp3_weight3_gs_dp1100_jack0.mat x 3553/0014_fgh_filtp3_weight3_gs_dp1100`
weighted rms 0.006 uK-deg over area 584.9 sq deg, QU combined: 0.006 uK-deg, 584.9 sq deg



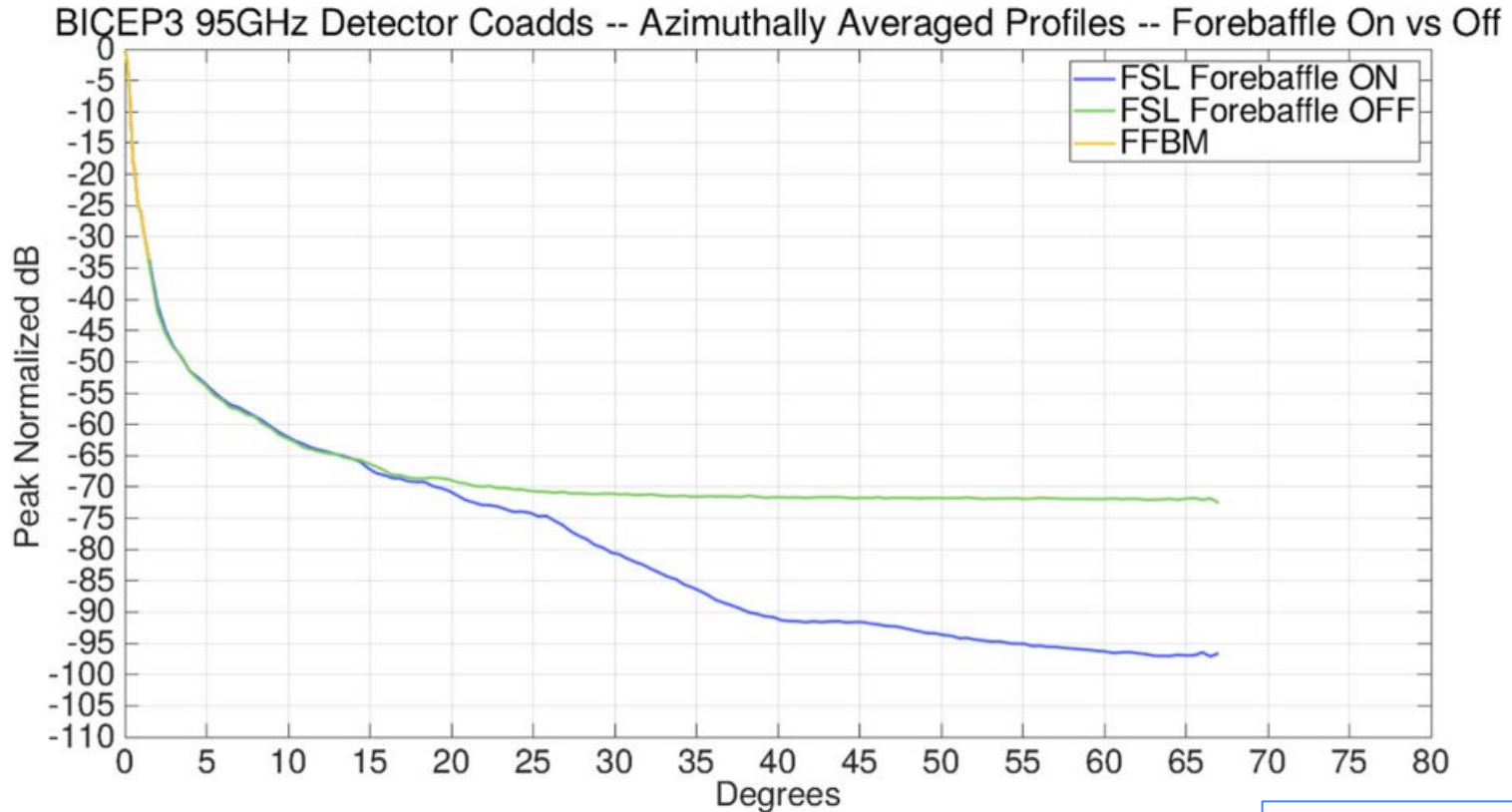
So where does the scan-sync signal come from? - why would it time evolve?

- ❖ In standard BICEP CMB scanning the azimuth range is held fixed for ~50 minutes while the sky drifts past (by ~12.5 degrees)
- ❖ We subtract the mean of all scans from each scan
 - This will perfectly remove any scan-sync signal which is constant in time
 - (done separately for left and right scans)
- ❖ Scan-sync signal is very large compared to the signal of interest
 - How time stable can we expect it to be? Where does it come from?

BICEP3 is very well shielded - but not perfectly!

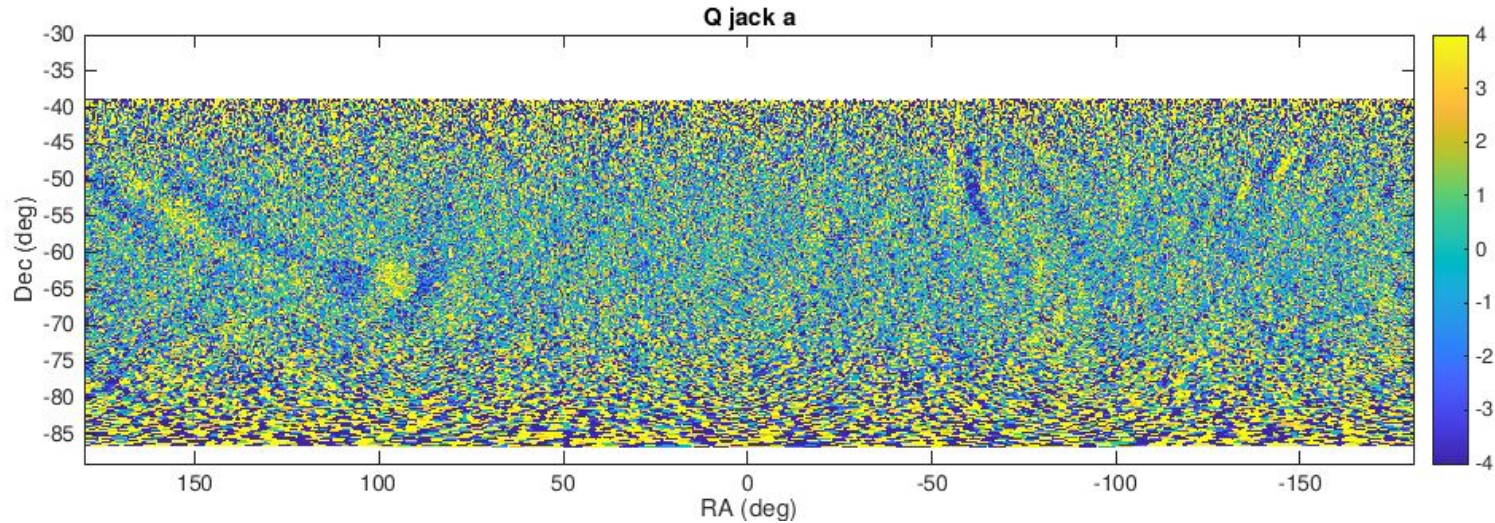


BICEP3 is very well shielded - but not perfectly!

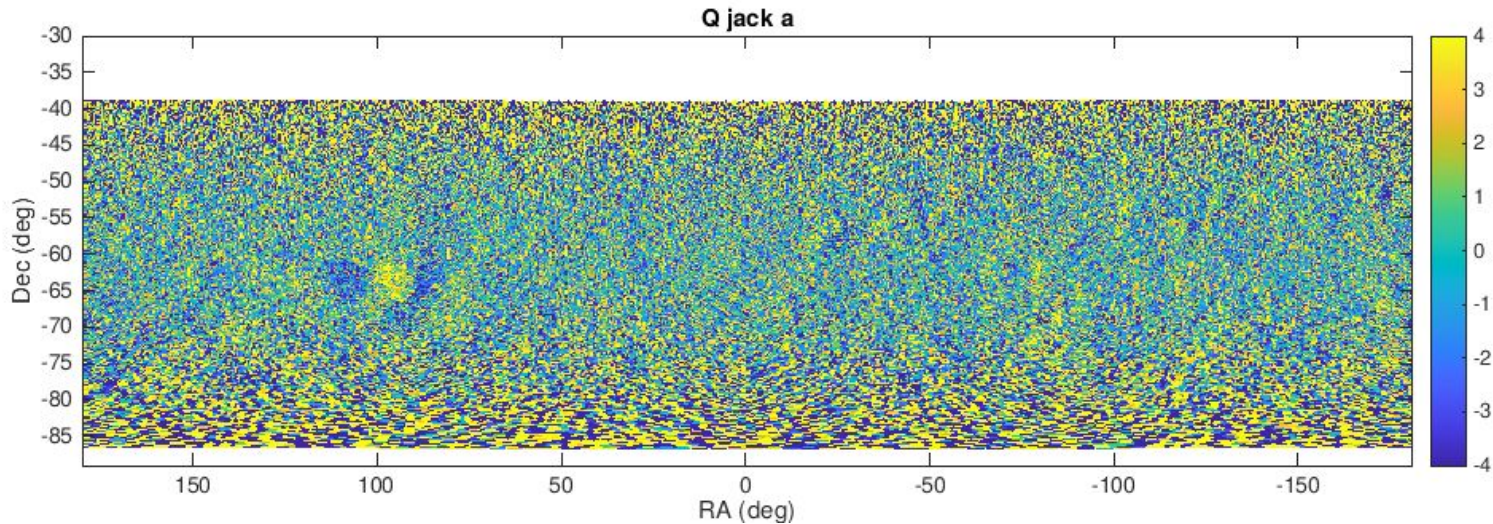


Unpublished work by
Christos Giannakopoulos

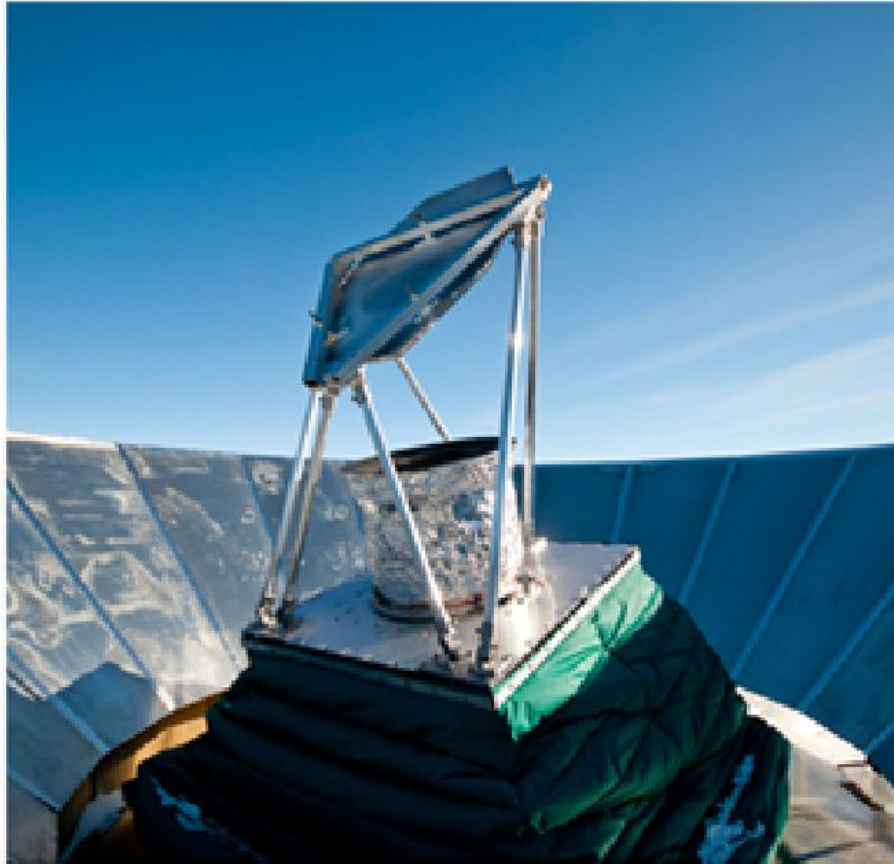
An example of a time evolving scan-sync signal: the Sun
Map: BICEP3 wide field, focal plane inner/outer split difference Q map
Case: Include all data



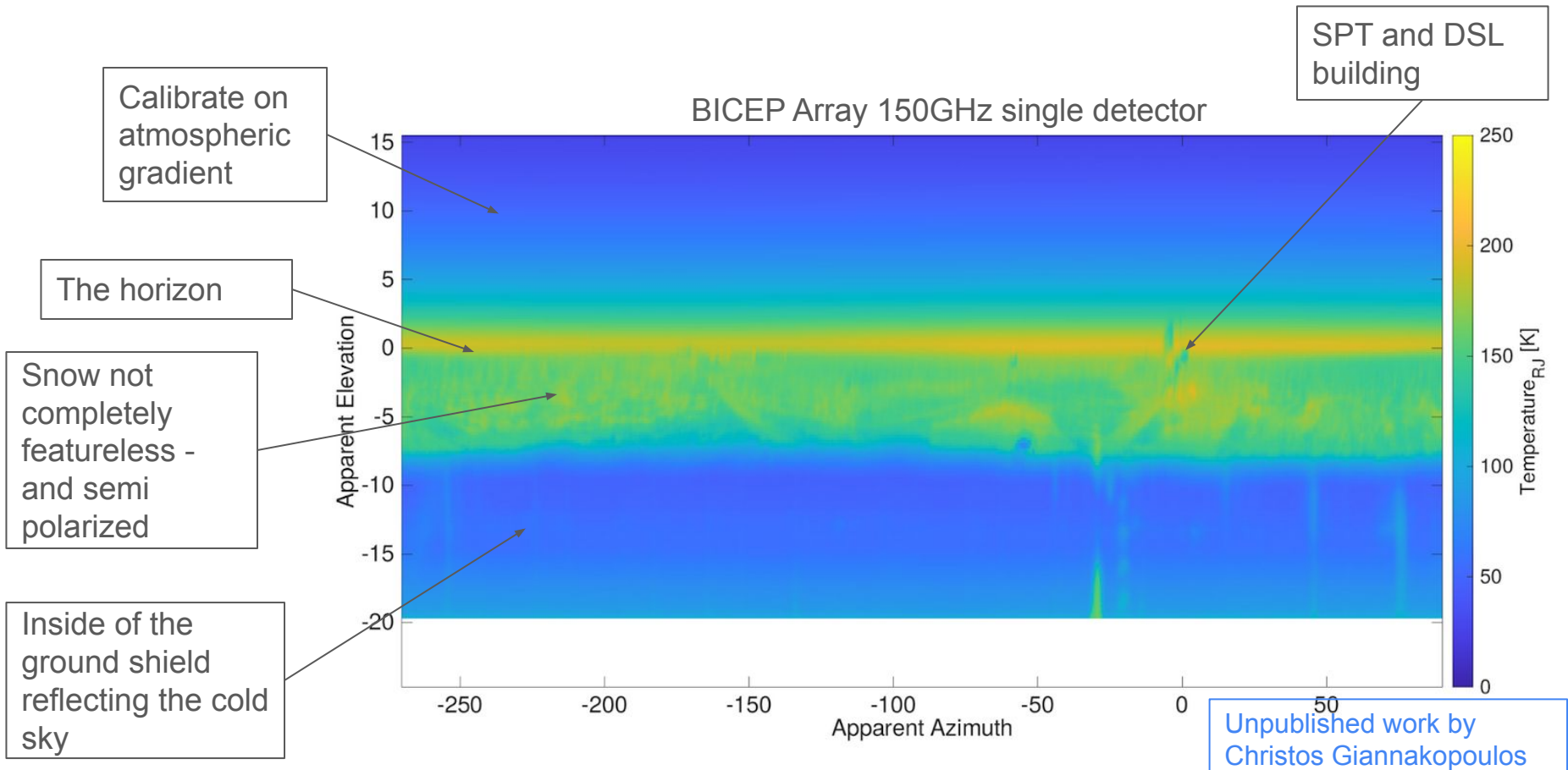
An example of a time evolving scan-sync signal: the Sun
Map: BICEP3 wide field, focal plane inner/outer split difference Q map
Case: Exclude when boresight <38 degrees from Sun



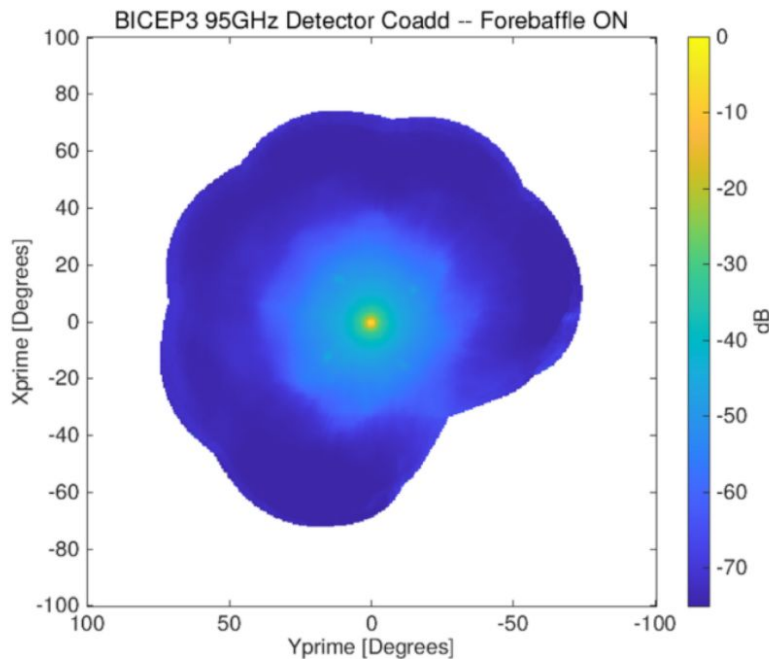
Put 45 degree flat mirror on and directly map the ground emission



Put 45 degree flat mirror on and directly map the ground emission



Try to predict scan-sync?! Convolve environment map with 4pi beam?



Unpublished work by
Christos Giannakopoulos

- ❖ Left is measured coadd - would need per-detector
- ❖ Full physical optics calculation (per detector) plus anechoic chamber measurements? (Planck did this)
- ❖ How to get environment map for Chile? SO could measure? S4 test telescope in Chile?
- ❖ How to add time-evolution of environment map? Again SO could measure. Could also try to model diurnal heating of mountains etc?

How to quantify systematics and risk therefrom?

- ❖ [John Ruhl](#) and [Colin Bischoff](#) showed yesterday that it is simple enough to make an “aggressive optics” all-Chile S4 config which nominally meets the charge from [Jim Strait’s presentation](#) (slide 9).
- ❖ There may be noise-like effects like polarized atmosphere which slow things down versus such existing forecasts...
- ❖ But how to quantify the risk that such a plan will never pass null tests - i.e. will contain demonstrable bias much larger than $r=0.0005$?
- ❖ It seems hard to argue that paper studies alone can ever reduce the risk to “acceptable” levels...
- ❖ It seems like we need a design development phase including actual hardware, both in the lab (antenna range tests) and in the field. This would be a big effort but still small compared to S4 total cost.

Discussion...

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- ❖ De-lensing validation
- ❖ Map based validation
- ❖ Optimizing band allocation on SATs and delensing LATs
- ❖ Inclusion of potential systematics

Questions:

- Path for survey optimization (sky coverage, scan strategy, LAT-SAT sensitivity balance)?
- More generally: Chilean alternatives, i.e. “A - forecast for a pre-defined Chilean survey” vs. “B - optimization and forecasts for a Chilean survey”?
- Can we do a suite of simulated systematics (simple, medium, high complexity) + get the mitigation / analysis pipeline ready in the allocated time? Or just have a plan?
- Can we trust a simulation? How to decide the level of systematics to put in?