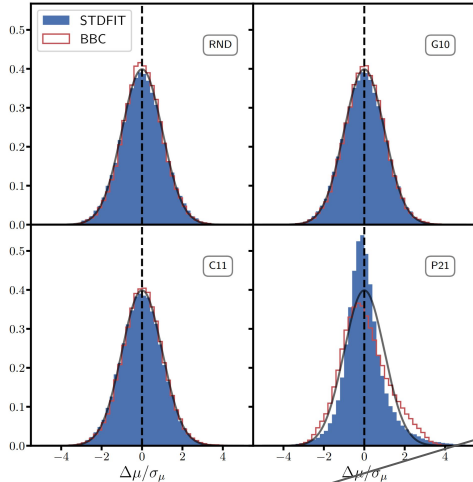


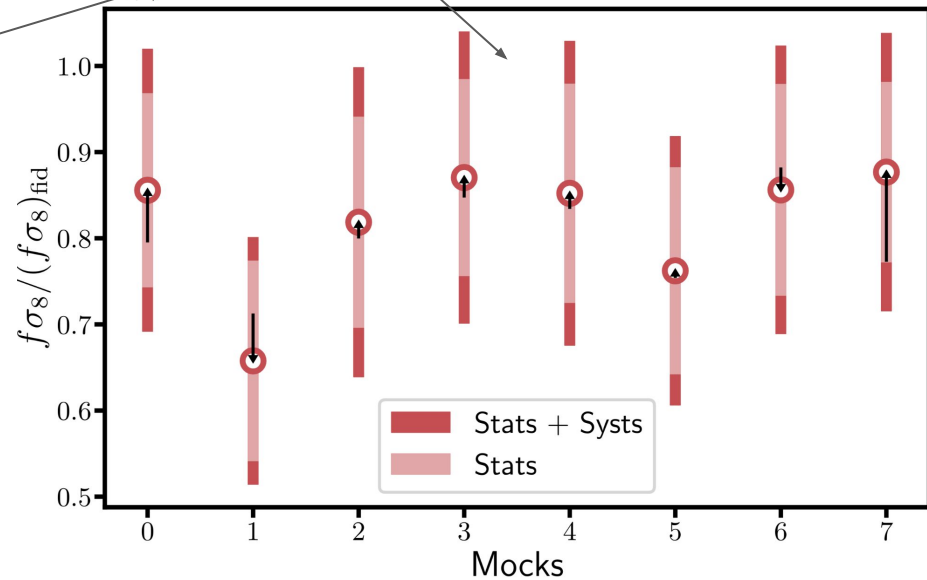
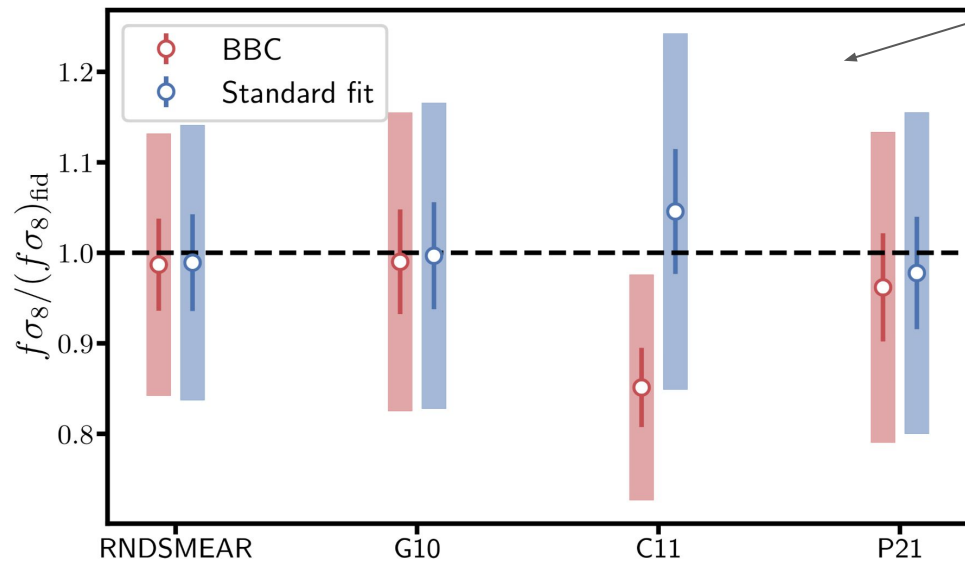
11/13/2024

Paper writing

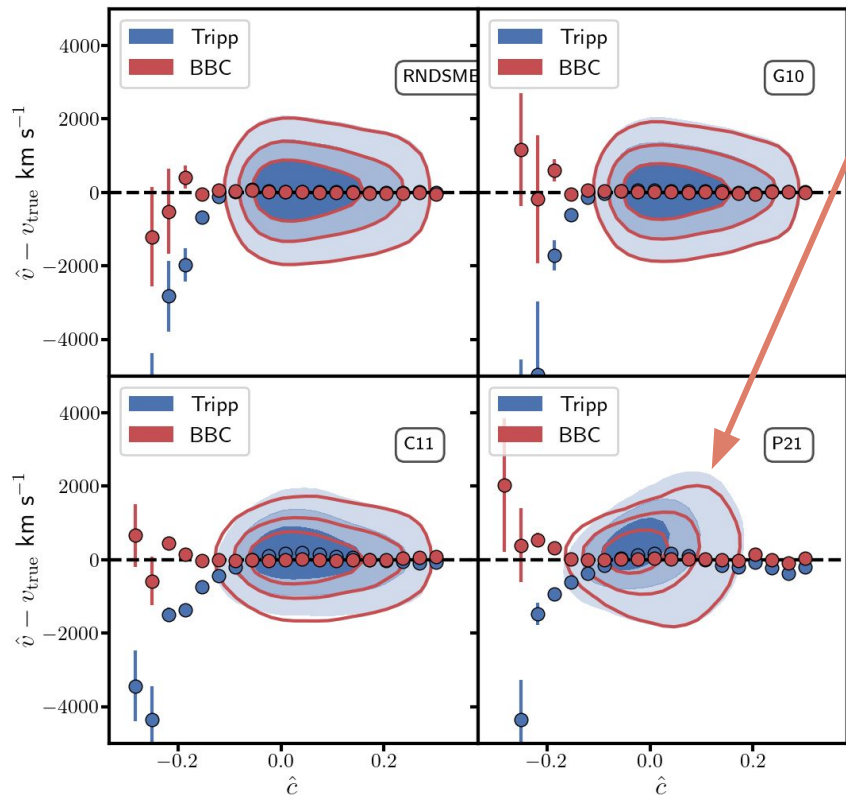
<https://www.overleaf.com/read/dvrvchwvrxhp#63525a>



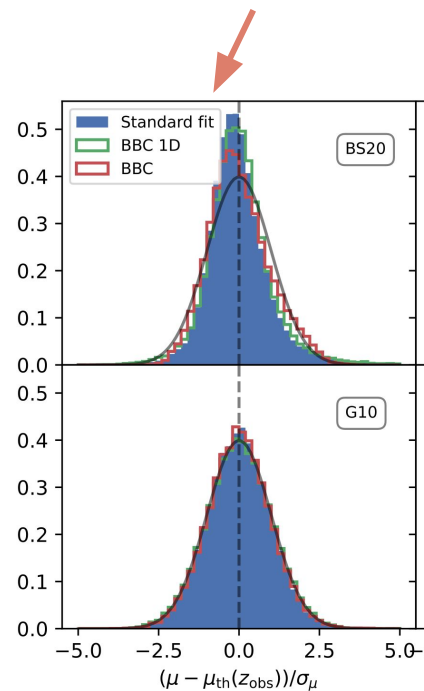
Main plots



10/16/2024



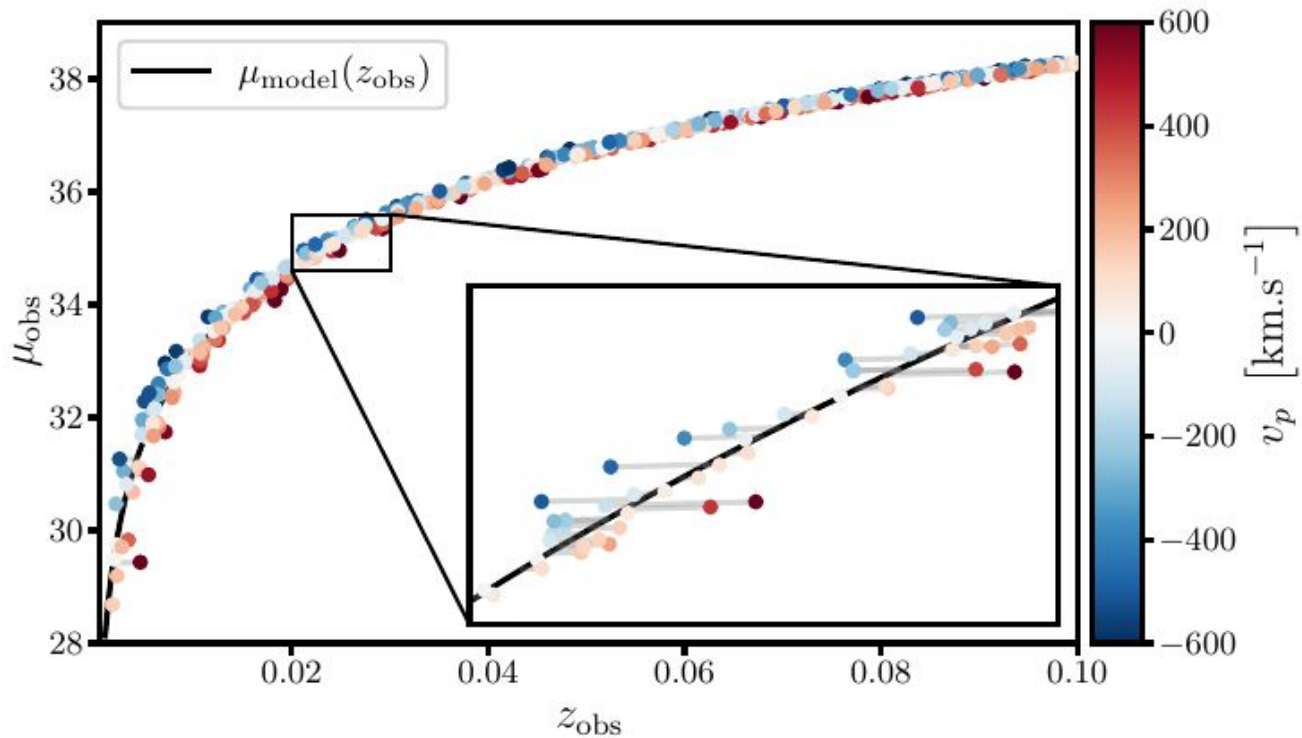
Bias correction working $\langle \mu \rangle = 0$ in c bin, but non gaussian distribution !



Few question for BBC experts:

- What does *bias on MU instead of $mB,x1,c$* means?
- What does use *avg mag from cosmology (not $nommag0$)* means?

PVs with SNe Ia : noise and probe



$f\sigma_8$ as a probe for general relativity

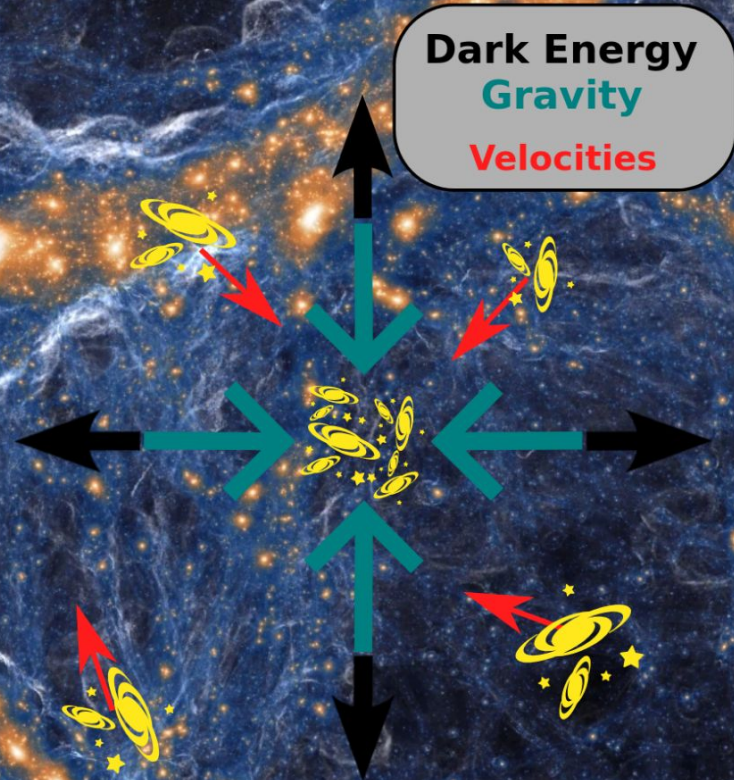
Velocities are linked to density through the continuity equation:

$$\nabla \cdot v(\mathbf{x}) \propto f\sigma_8 \tilde{\delta}(\mathbf{x})$$

where $f \equiv$ growth rate

General Relativity + Λ CDM:

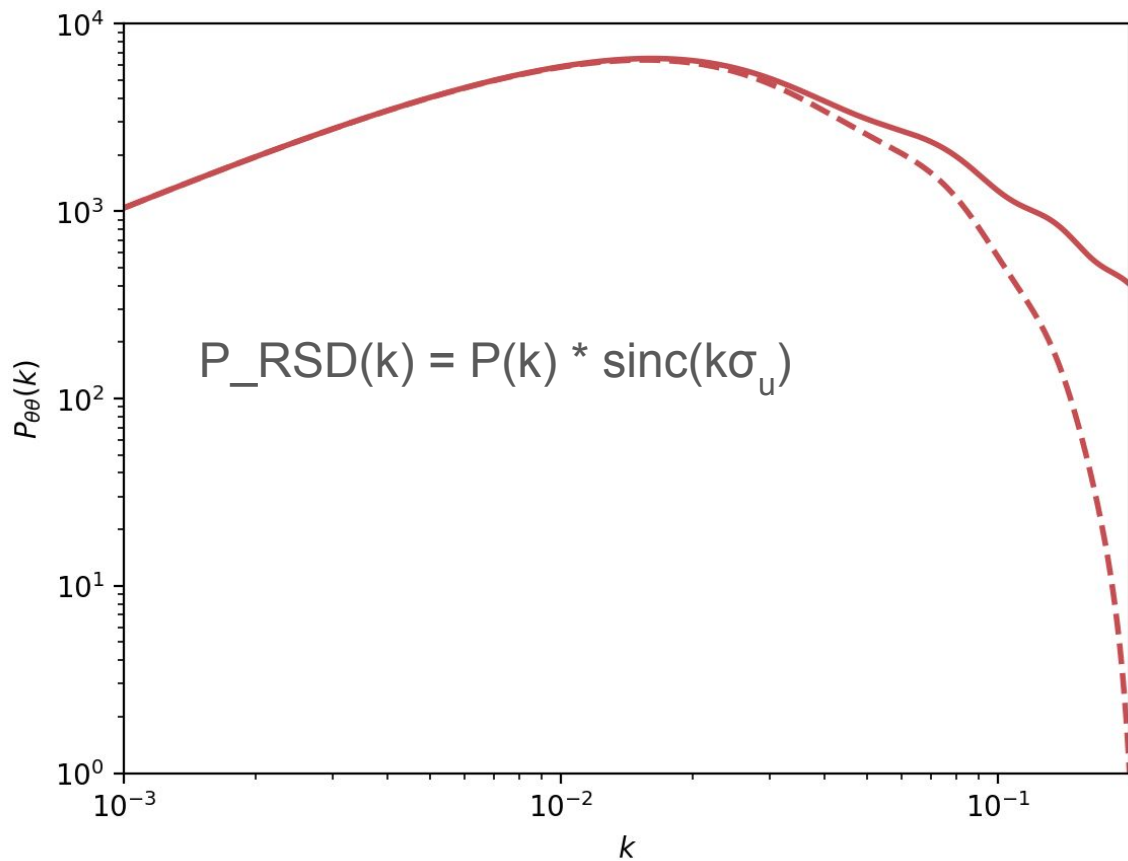
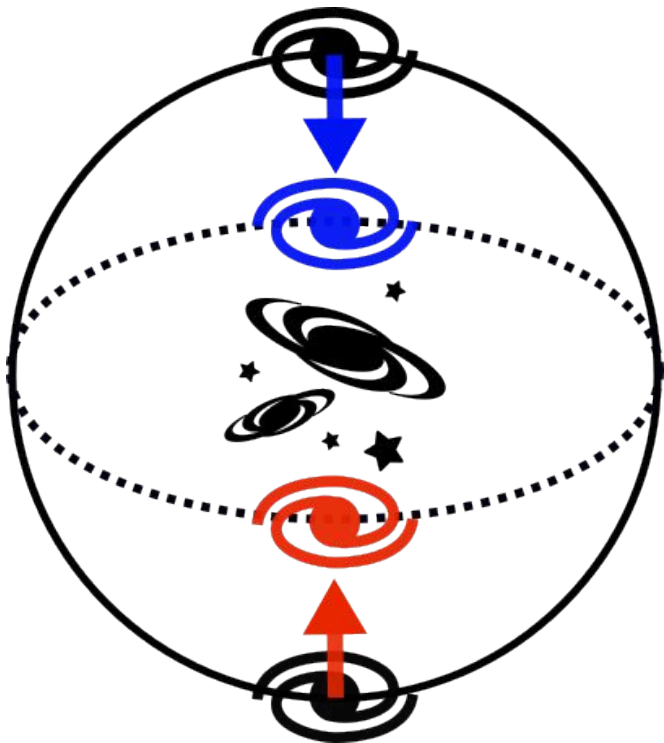
$$f \simeq \Omega_m^\gamma \text{ with } \gamma \simeq 0.55$$



- Writing the growth-rate with BBC paper
- Fitting σ_u (the thorn in the side of growth-rate measurement) for Uchuu

06/26/2024

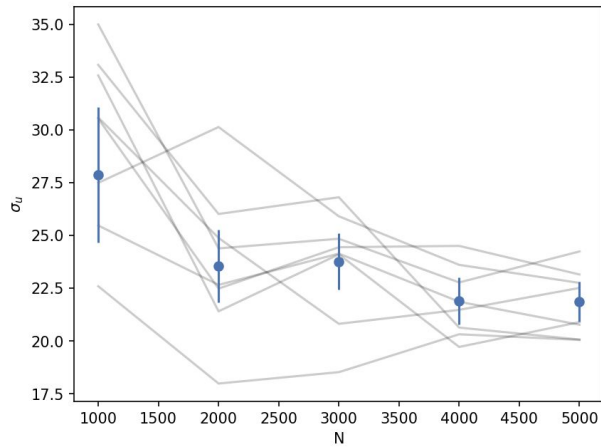
Week Recap



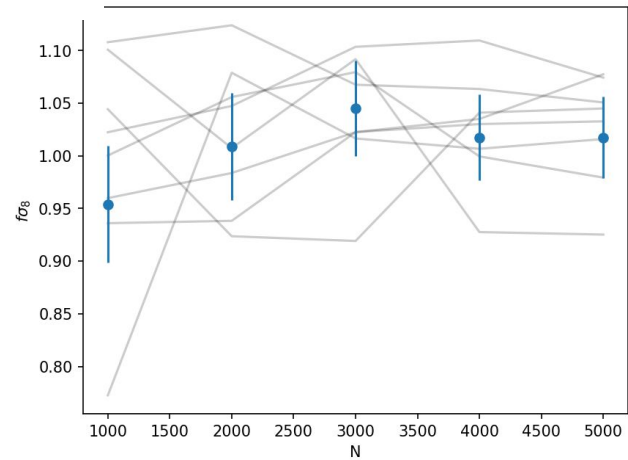
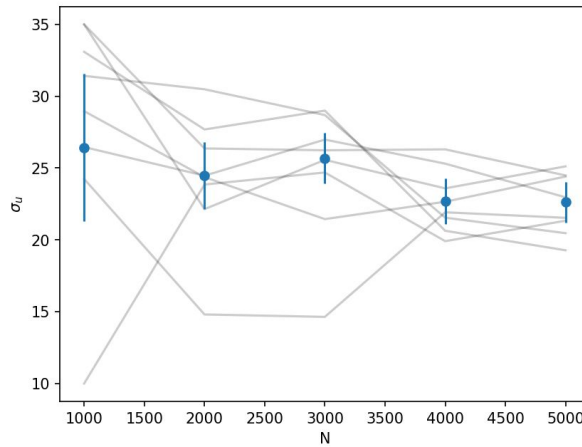
06/26/2024

Week Recap

fs8 fixed



fs8 free

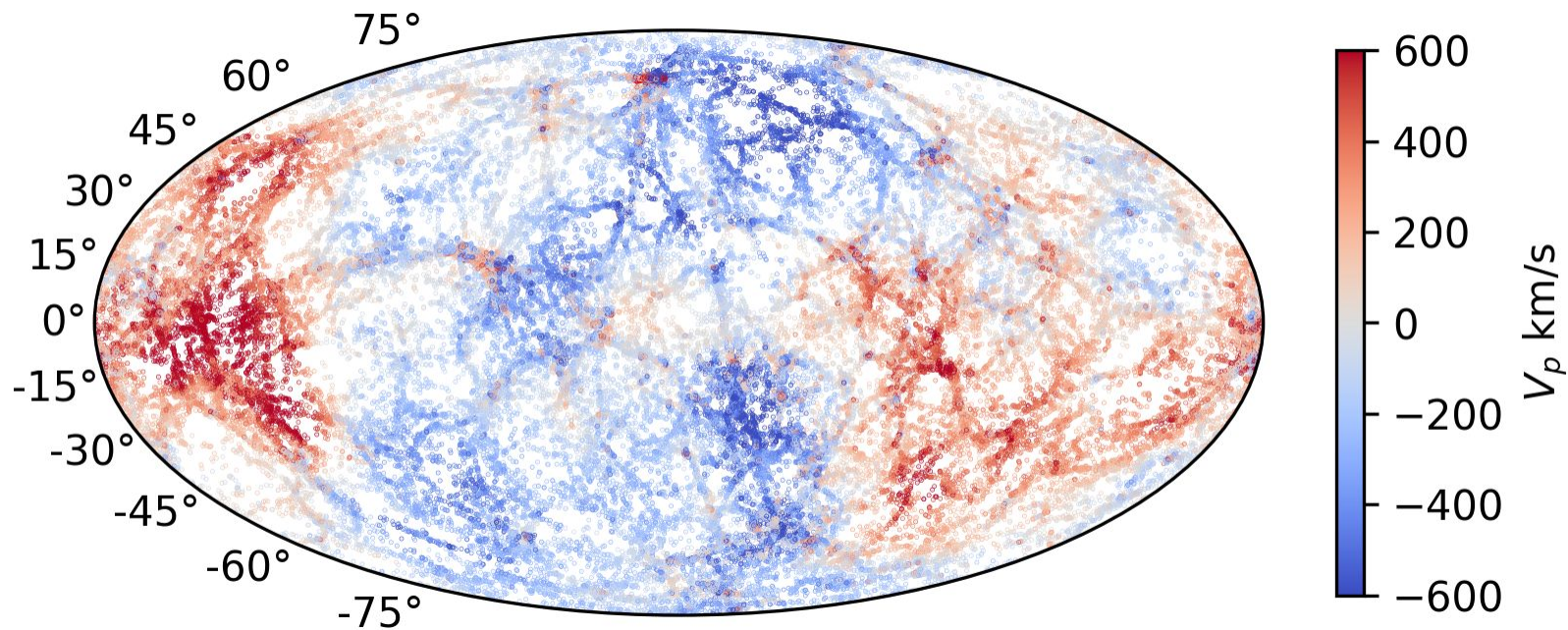


- Code to apply selection to an existing SNANA sim
- Start writing simulation section for UGGs
- Have a look to correlation function of velocities in Uchuu

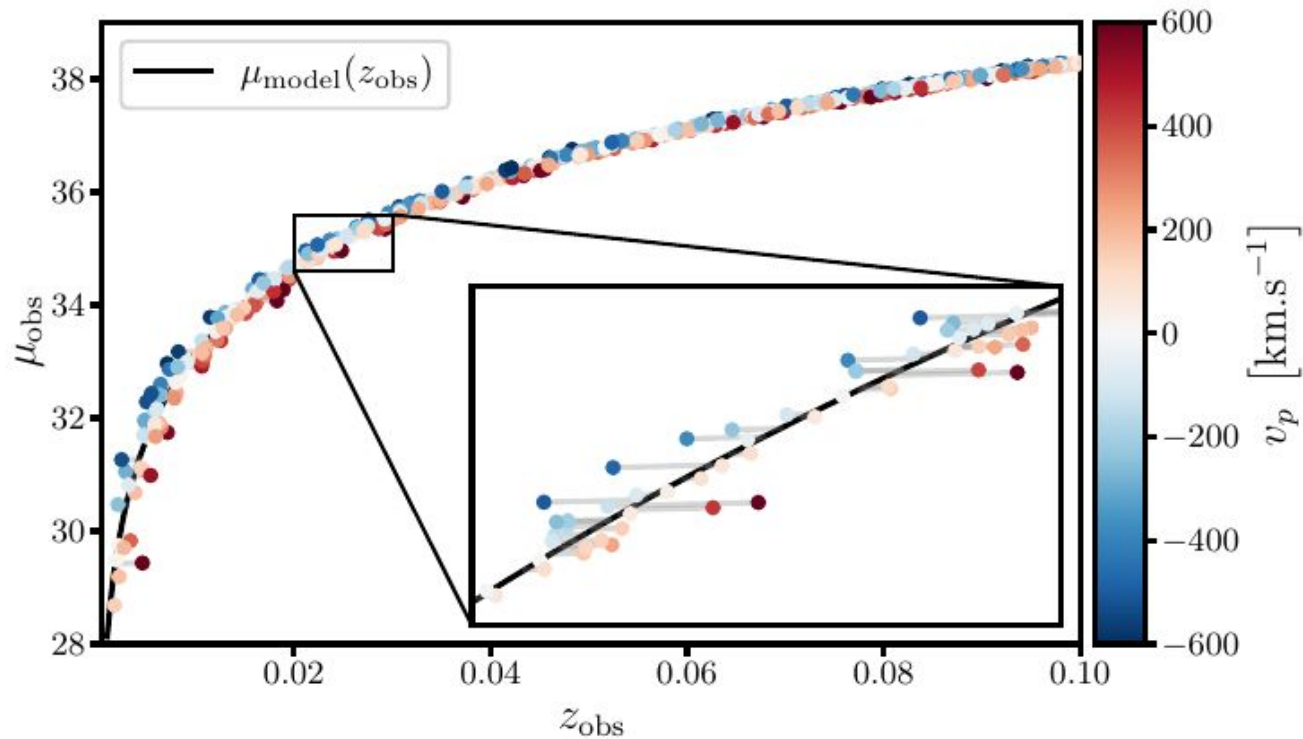
06/26/2024

Week Recap

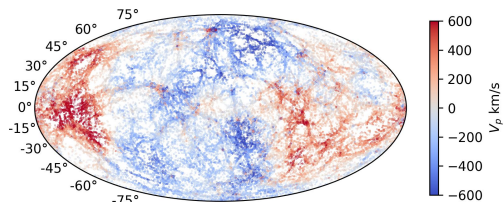
The PV field is correlated on large scales...



....And affect the Hubble Diagram of SNe Ia...

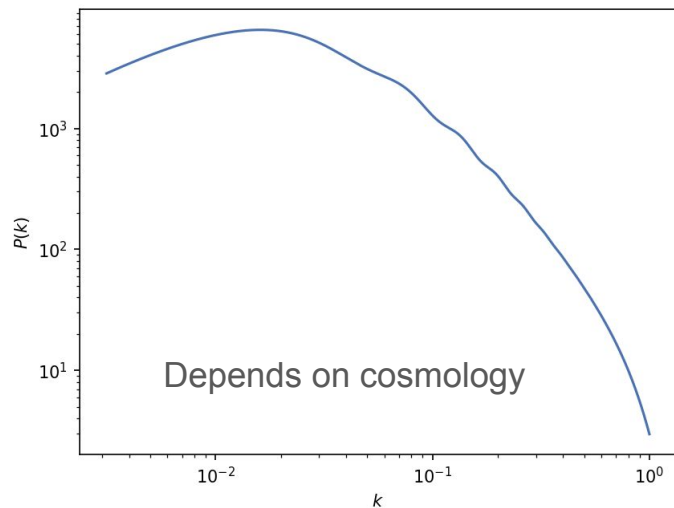


...but you can use this correlation to get cosmological constraints



$$C_{ij}^{vv} = \frac{H_0^2 f^2}{2\pi^2} \int dk P_{\theta\theta}(k) W(k; \mathbf{x}_i, \mathbf{x}_j)$$

Geometrical
stuff



For projection of PVs // and perp to
separation vector \mathbf{r} :

$$\Psi_{\perp, \parallel}(r) = \frac{H_0^2 \Omega_0^{1.2}}{2\pi^2} \int_0^\infty P(k) K_{\perp, \parallel}(kr) dk, \quad (2)$$

where

$$K_{\perp}(x) = \frac{j_1(x)}{x}, \quad K_{\parallel}(x) = j_0(x) - 2 \frac{j_1(x)}{x},$$

and

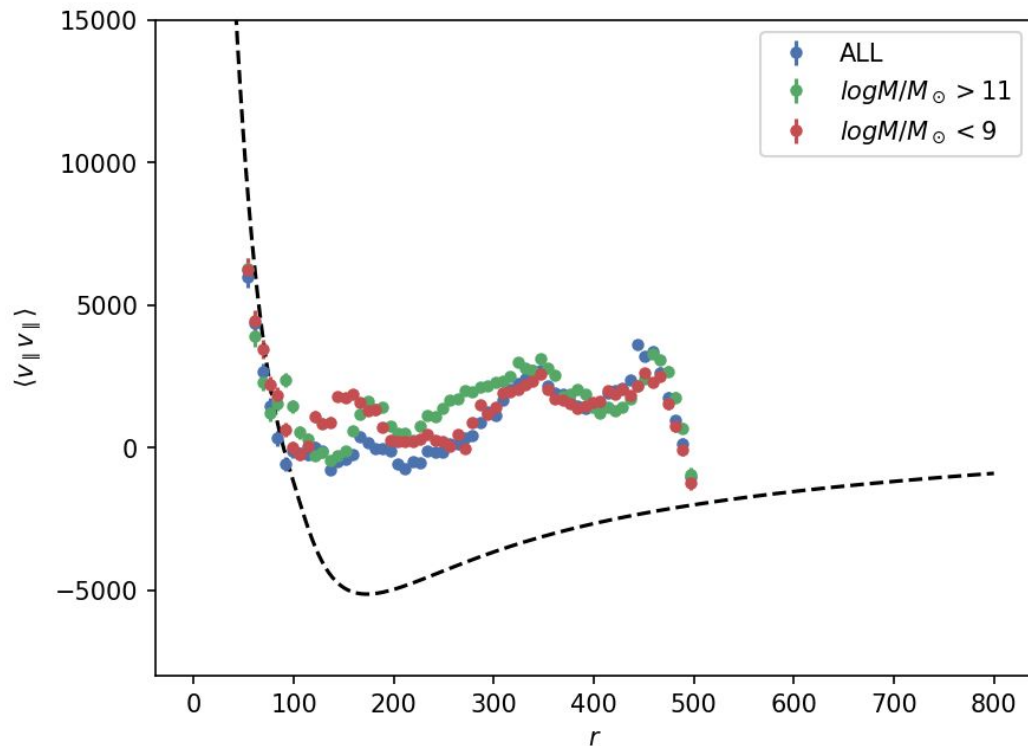
$$j_0(x) = \frac{\sin x}{x}, \quad j_1(x) = \frac{\sin x}{x^2} - \frac{\cos x}{x}.$$

$$\langle v_{1,\parallel}(x)v_{2,\parallel}(x+r) \rangle = \frac{1}{N} \sum_i v_{1,\parallel} v_{2,\parallel}$$

Is the 2pts statistics different for galaxies of different masses ?

Good point: don't seem to depend on mass

Bad point: don't fit the theoretical curve

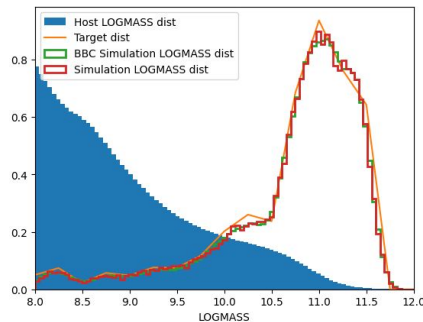
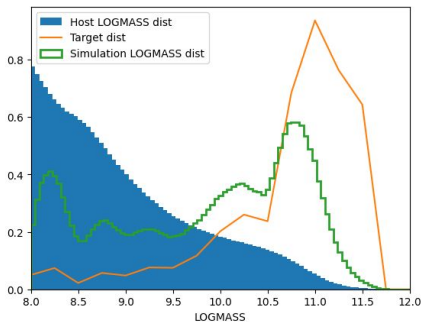


05/28/2024

Week Recap

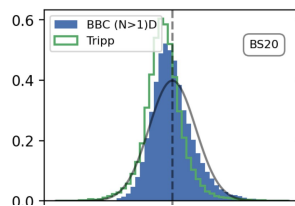


- Work on FS8 / BBC:
 - Solve a host distribution problem

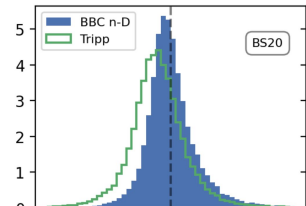


- Solve minor problems inside pippin file (e.g. MW model in sim and fit conf were inconsistent) and reorganize for simulation in multiple mocks => 8 Mocks = 7 “data” mocks + 1 BBC sim mock

- BS20 Problem still here (is it really a problem or should it be like this?)



HD Pull



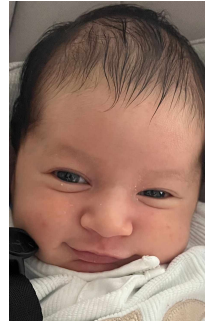
HD Residuals

- Got back my ZTF paper referee report: No major fix to do!



04/30/2024

Week Recap



I'll never
let you
sleep

- Still stuck with the BS20 model problem... (NERSC is in maintenance today)
- New version of the simplified FoF that run faster => currently running on full mock. The goal is to find the expected vel scattering inside group, wrt mass, number of gal, etc... then compare with Erik's data.
- Trying to fit fs8 with current simulations: developing and debugging of (Field Level Inference Package)



Quick SNANA question:

In fitfiles some HOST_RA, DEC are set to -999 however from the hostlib these RA, DEC are correctly defined (even MW seems correctly computed for them). Did someone already encounter this behaviour?

04/17/2024

Week Recap

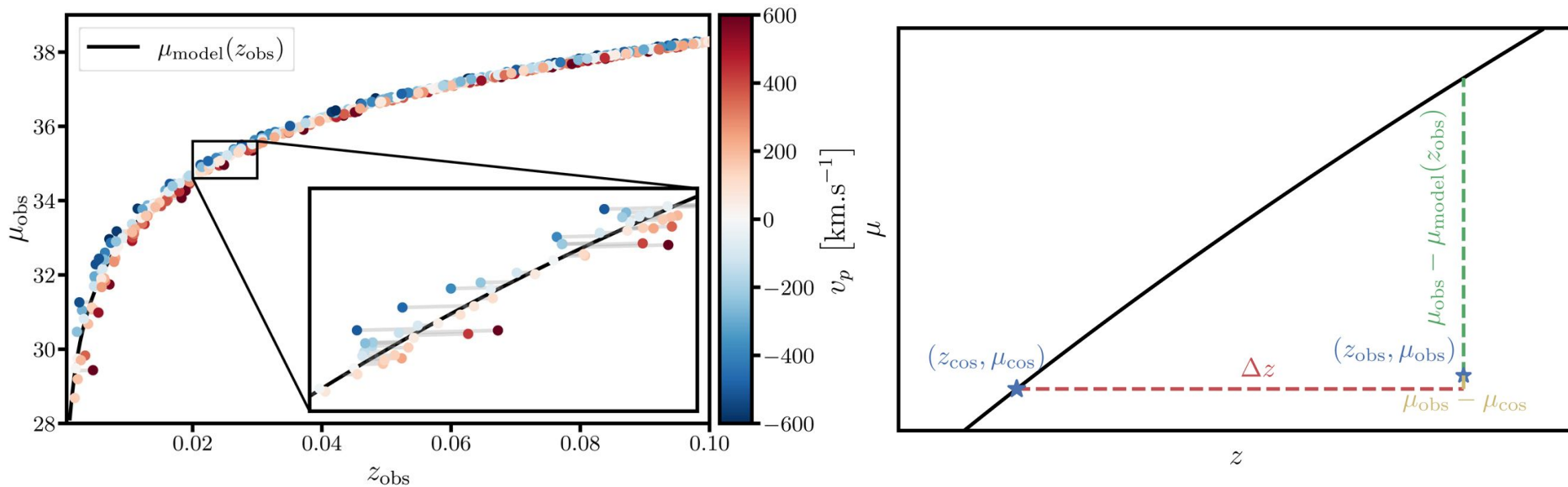
- Taking care of this cute little baby



- Get back to work on pec. vel. project

04/17/2024

Project: fsigma8 with LSST survey - BBC framework



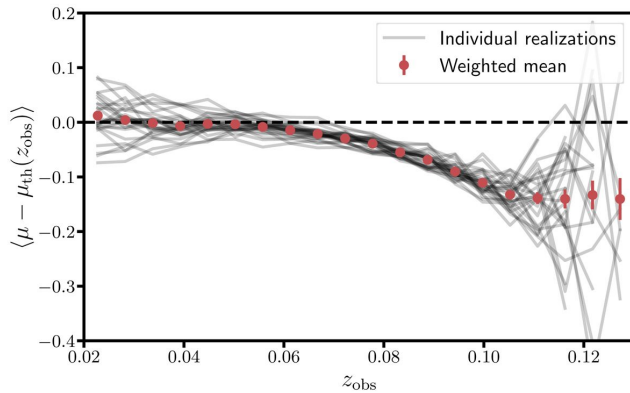
$$\hat{v}_i(\mathbf{p}_{\text{HD}}) = -\frac{\ln(10)c}{5} \left(\frac{(1+z_i)c}{H(z_i)r(z_i)} - 1 \right)^{-1} \Delta\mu_i(\mathbf{p}_{\text{HD}}).$$

04/17/2024

Until now: only sim that assume color indep-scattering but color dep scattering brings : asymmetry in residuals, bias in std parameters, ...

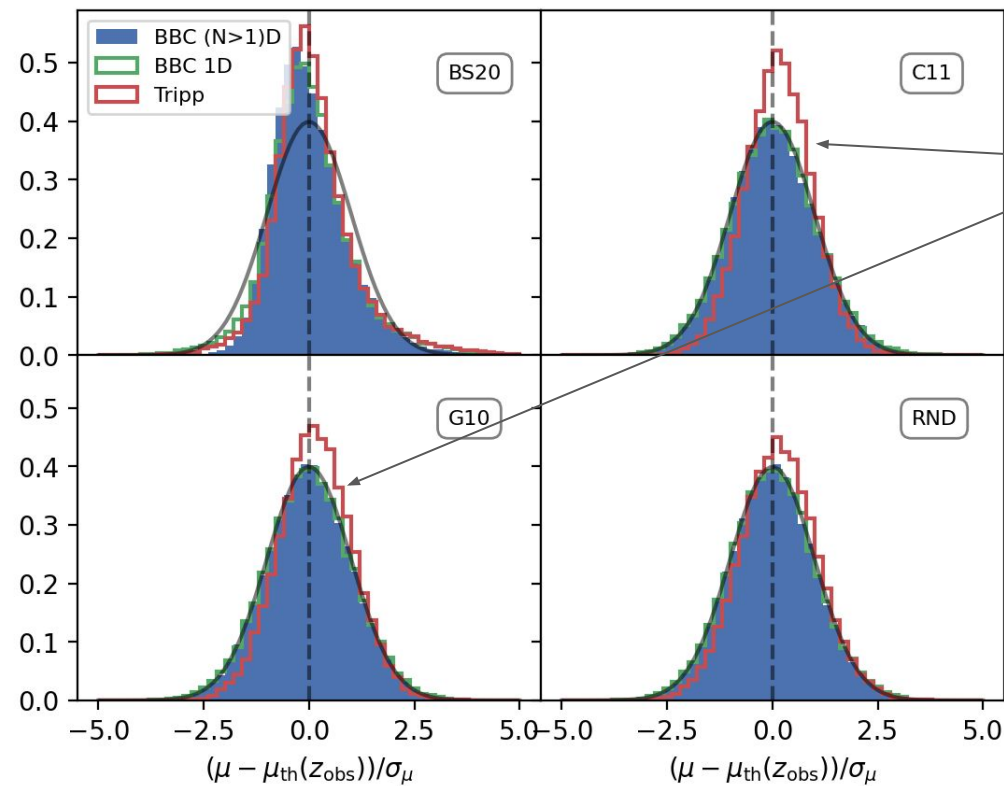
Primary Goal: see if the BBC framework help to correct these bias/compare with std method

Secondary goal: If we don't have a complete sample due to mag-limited selection (typing/spec-z/photometry) could bbc correct for that without biasing pec. vel.?



04/17/2024

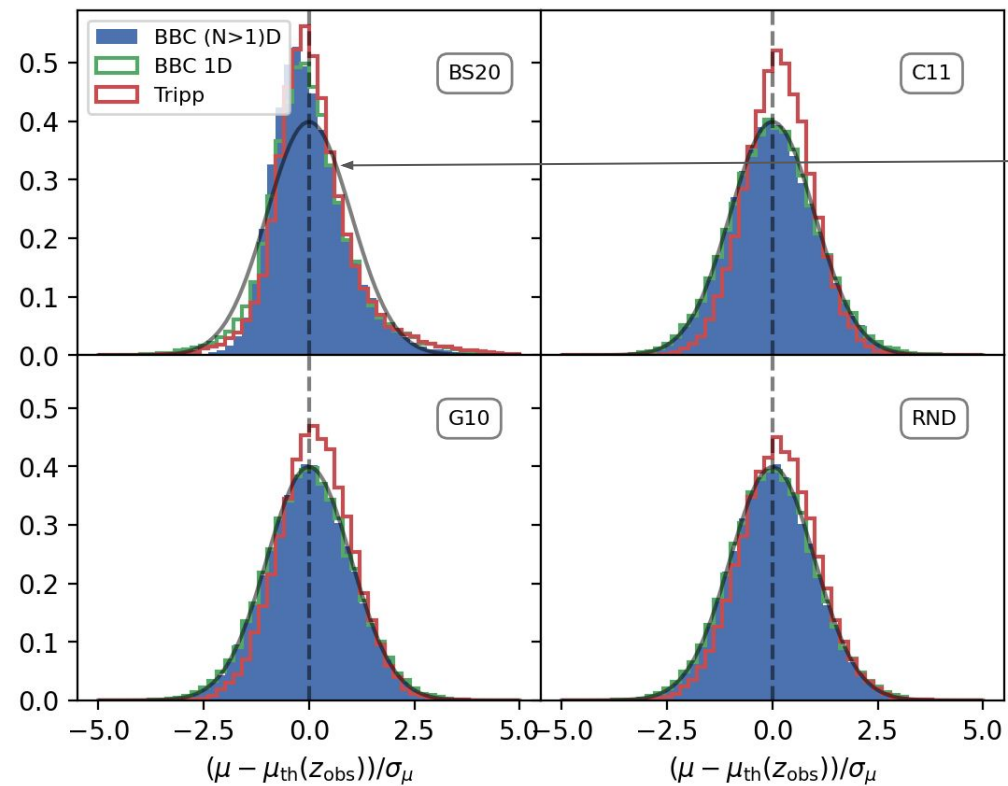
HD Pull



Bias correction seems to give better pull than standard tripp

04/17/2024

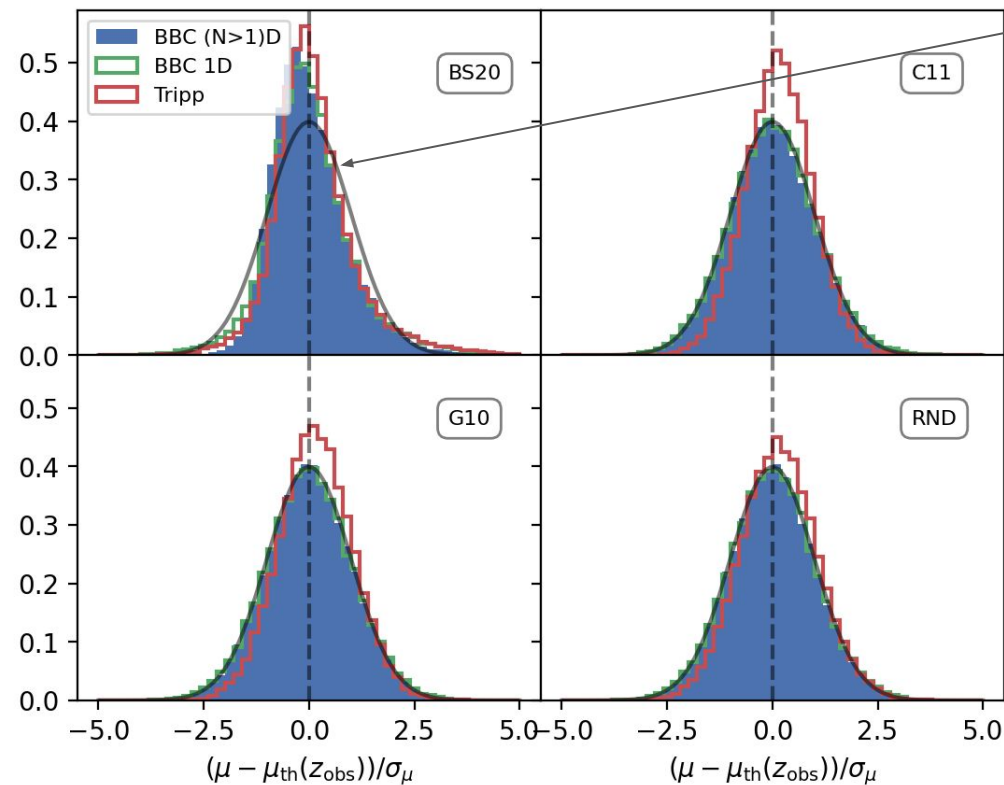
HD Pull



Problem with the 4D bias correction:
Currently trying to look at other config
used on midway

04/17/2024

HD Pull



Problem with the 4D bias correction:
Currently trying to look at other config
used on midway

Bias seems color dependant and
does not correspond to a simple error
offset

