

ALMA-IMF Data Reduction Telecon, June 5, 2020

Time: 0900 EDT / 1300 UTC June 5, 2020

Unirse a la reunión Zoom

<https://cuaed-unam.zoom.us/j/91484486311>

ID de reunión: 914 8448 6311

Contraseña (password): dr_meeting

[Last week's notes](#)

[Next week's notes](#)

Attendees: Adam, Roberto, manuel, Tapas, Fred, Fab, Patricio, Mélisse, Thomas, Estrella, Amy, Hongli, Hongli, Yohan

Not attending:

Agenda:

- Noise Calculation Discussion
 - Adam's talk in the science workshop shows that desired noise is rarely reached. Adam requested us to check if they are appropriate and go back to him.
https://github.com/ALMA-IMF/reduction/tree/master/reduction/noise_estimation_regions
- Problems that will affect line intensity (and maybe continuum):
 - <https://help.almascience.org/index.php?na/Knowledgebase/Article/View/419>
(Ana notified us about this)
 - This tool:
https://github.com/ALMA-IMF/reduction/blob/master/reduction/plot_autocorrelations.py can be used to plot autocorrelations to see if a bright line affects the data
 - In many cases, the CO 2-1 line is clearly evident in the autocorr spectrumPatricio: also the weblog Tsys plots can be used to see this, and maybe we even saw it when we started reduction. It may affect more the 7M data.
Timea: Have we already done these corrections?
Roberto: I think I made some tests of re-running the pipeline flagging out some Tsys ranges, with no obvious improvement. Will check.
https://docs.google.com/spreadsheets/d/1xM8AfiMpe8SVifqzI0IzjD10CEB2OILlw_Kw5R1sxAM/edit#gid=1692076735
- Full (CASA) cubes:
 - /bio/web/secure/adamginsburg/ALMA-IMF/fullcubes
 - Only temporarily at that location; for future access, need Globus
 - Roberto: i think we need to agree on two things:
 - 1) a set of default parameters to make these full spw cubes. My current take is concatenated visibilities and perchanweightdensity=False to have almost same

beam as continuum image, small value of chanchunks and hogbom deconvolver to to make it faster, "short-baseline" (name from CASA tutorial) auto-MT and deep cleaning (niter of millions for a full spw with 1920 chans, threshold at ~ 5 to 10 mJy). I'm still investigating the details of CVEL-ing visibilities prior to concat.

Details in:

<https://docs.google.com/document/d/1nNmpEcD-axjVrHRowfcK3clmaZ0xrapBRcWnKg-HkVI/edit#heading=h.wc0r6ja4cl6f>

Roberto will modify defaults in `imaging_parameters.py` and submit pull request

HongLi: maybe more complicated lines like CO need to set a restrictive `pblimit`.

Manuel: `pblimit = 0`, `pbmask = 0.2` to avoid convergence problems at edges

2) Do we make these cubes in a distributed way or can the Florida supercomputer handle all of them?

- IMAGER vs CASA
 - Are there benchmarks? Is IMAGER really better & faster?
 - If it is faster and better, how can we start using it?
Tests are being done both for continuum and cubes. Full spw cubes can be done in ~ 1 hour
- Cube beam sizes are often wrong
 - Problem seems to be `tclean(vis=[one, two, three, ...])` Can be resolved by cleaning a concatenated data set. However, we may be able to hack around this (and avoid unnecessary data duplication) by modifying the headers
 - Or we can just bite the bullet and accept yet another data duplication
- <https://bio.rc.ufl.edu/secure/adamginsburg/ALMA-IMF/May2020/quicklooks/>
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