

# UNIT 6: EVOLUTION OF STARS

## BINARY STAR MASSES

THANK YOU MRS. RAMSEYER!!! =)



Here are two Nice Reference Websites:

- [The Masses of Stars](#) (copyrighted)
- [Eclipsing Binary Simulator](#) **REALLY AWESOME!!**

1. If two stars orbit each other with a period of 10 years and a separation of 5 AU, what is their total mass?

$$r = 5 \text{ AU} \qquad P = 10 \text{ Years} \qquad M = (m_A + m_B) = \text{Total Mass} = ?$$

$$M = (m_A + m_B) = \frac{r^3}{P^2} \qquad M_T = m_a + m_b = \frac{5^3}{10^2} = 1.25 \text{ Solar Masses} = 1.25 M_\odot$$

$M_\odot$  is the symbol for the mass of the Sun,  $1.989 \times 10^{30}$  kg

2. What is the orbital period of a bit of matter in an accretion disk  $2 \times 10^5$  km (convert to AU) from a 10 solar mass black hole?

$$r = 2 \times 10^5 \text{ km} \qquad P = ? \qquad M = (m_A + m_B) = \text{Total Mass} = 10 M_\odot$$

$$r = \frac{2 \times 10^5 \text{ km}}{1} \times \frac{1 \text{ AU}}{1.5 \times 10^8 \text{ km}} = 0.0013 \text{ AU}$$

$$M = (m_A + m_B) = \frac{r^3}{P^2}$$

$$P^2 = \frac{r^3}{M}$$

$$P = \sqrt{\frac{r^3}{M}} = \sqrt{\frac{(0.0013)^3}{10}} = 1.5 \times 10^{-5} \text{ years}$$

$$\frac{1.5 \times 10^{-5} \text{ years}}{1} \times \frac{365 \text{ days}}{1 \text{ year}} \times \frac{24 \text{ hr}}{1 \text{ day}} \times \frac{60 \text{ min}}{1 \text{ hr}} = 8 \text{ minutes}$$

3. If SS 433 consists of a 20 solar mass star and a neutron star (3 solar masses) orbiting each other every 13.1 days (convert to years), then what is the average distance between them?

$$r = ? \qquad P = 13.1 \text{ days} \qquad M = (m_A + m_B) = 20 M_\odot + 3 M_\odot = 23 M_\odot$$

$$r = \frac{13.1 \text{ days}}{1} \times \frac{1 \text{ year}}{365 \text{ days}} = 0.0359 \text{ years}$$

$$r^3 = P^2 M \qquad r = \sqrt[3]{MP^2} = \sqrt[3]{(0.036)^2 (23)} = 0.31 \text{ AU}$$

4. If Star A, a 1 solar mass star, is 2 AU from the center of mass of a binary system and Star B is 42 AU distant, what is the mass of Star B?

$$m_A = 1 M_{\odot} \quad r_A = 2 \text{ AU} \quad m_B = ? \quad r_B = 42 \text{ AU}$$

$$\frac{m_A}{m_B} = \frac{r_B}{r_A} \quad \frac{1 M_{\odot}}{m_B} = \frac{42 \text{ AU}}{2 \text{ AU}} \quad m_B = 0.048 M_{\odot}$$

5. Star A has a mass of 10 solar masses and wobbles about 2 AU from an apparent center of mass. An apparent X-Ray source exists 4 AU from this center of mass. How massive is this dark companion?

$$m_A = 10 M_{\odot} \quad r_A = 2 \text{ AU} \quad m_B = ? \quad r_B = 4 \text{ AU}$$

$$\frac{m_A}{m_B} = \frac{r_B}{r_A} \quad \frac{10 M_{\odot}}{m_B} = \frac{4 \text{ AU}}{2 \text{ AU}} \quad m_B = 5 M_{\odot}$$